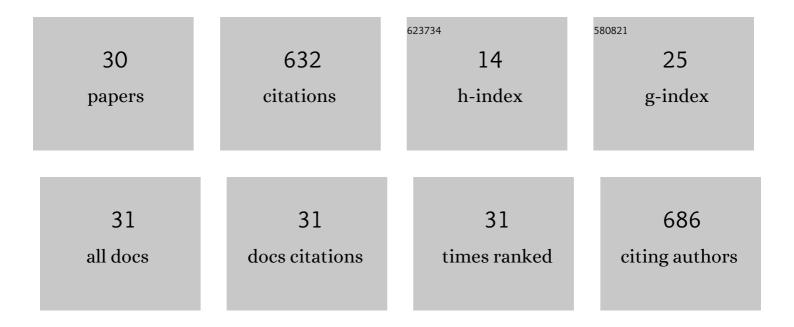
Peter Gomory

List of Publications by Year in descending order

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#	Article	IF	CITATIONS
1	Acceleration in Fast Halo CMEs and Synchronized Flare HXR Bursts. Astrophysical Journal, 2008, 673, L95-L98.	4.5	173
2	SOLAR MAGNETIZED TORNADOES: ROTATIONAL MOTION IN A TORNADO-LIKE PROMINENCE. Astrophysical Journal Letters, 2014, 785, L2.	8.3	49
3	Magnetic loop emergence within a granule. Astronomy and Astrophysics, 2010, 511, A14.	5.1	48
4	PLASMA DIAGNOSTICS OF AN EIT WAVE OBSERVED BY <i>HINODE</i> /EIS AND <i>SDO</i> /AIA. Astrophysical Journal Letters, 2011, 743, L10.	8.3	43
5	MULTIWAVELENGTH IMAGING AND SPECTROSCOPY OF CHROMOSPHERIC EVAPORATION IN AN M-CLASS SOLAR FLARE. Astrophysical Journal, 2010, 719, 655-670.	4.5	36
6	SPECTROSCOPIC OBSERVATIONS OF A CORONAL MORETON WAVE. Astrophysical Journal Letters, 2011, 737, L4.	8.3	36
7	Spectroscopy and Differential Emission Measure Diagnostics of a Coronal Dimming Associated with a Fast Halo CME. Astrophysical Journal, 2019, 879, 85.	4.5	28
8	Generation Mechanisms of Quasi-parallel and Quasi-circular Flare Ribbons in a Confined Flare. Astrophysical Journal, 2017, 847, 124.	4.5	26
9	Observations of a Footpoint Drift of an Erupting Flux Rope. Astrophysical Journal, 2019, 883, 96.	4.5	21
10	The three-dimensional structure of sunspots. Astronomy and Astrophysics, 2008, 488, 1085-1092.	5.1	21
11	Multi-wavelength fine structure and mass flows in solar microflares. Astronomy and Astrophysics, 2009, 505, 811-823.	5.1	19
12	Properties of the inner penumbral boundary and temporal evolution of a decaying sunspot. Astronomy and Astrophysics, 2018, 620, A191.	5.1	17
13	The Solar Activity Monitor Network – SAMNet. Journal of Space Weather and Space Climate, 2022, 12, 2.	3.3	16
14	Pre-eruption Processes: Heating, Particle Acceleration, and the Formation of a Hot Channel before the 2012 October 20 M9.0 Limb Flare. Astrophysical Journal, 2019, 874, 122.	4.5	15
15	Temporal evolution of arch filaments as seen in He†I 10 830 à Astronomy and Astrophysics, 2018, 617, A55.	5.1	14
16	Fitting peculiar spectral profiles in He <scp>I</scp> 10830 Ã absorption features. Astronomische Nachrichten, 2016, 337, 1057-1063.	1.2	12
17	Broadband microwave sub-second pulsations in an expanding coronal loop of the 2011 August 10 flare. Astronomy and Astrophysics, 2016, 593, A80.	5.1	9
18	Chromospheric evaporation flows and density changes deduced from Hinode/EIS during an M1.6 flare. Astronomy and Astrophysics, 2016, 588, A6.	5.1	9

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19	Spectropolarimetric observations of an arch filament system with the GREGOR solar telescope. Astronomische Nachrichten, 2016, 337, 1050-1056.	1.2	9
20	SOHO/CDS observations of waves above the network. Astronomy and Astrophysics, 2006, 448, 1169-1175.	5.1	9
21	Flare-induced changes of the photospheric magnetic field in a <i>δ</i> -spot deduced from ground-based observations. Astronomy and Astrophysics, 2017, 602, A60.	5.1	6
22	Evidence of quiet-Sun chromospheric activity related to an emerging small-scale magnetic loop. Astronomy and Astrophysics, 2013, 556, A7.	5.1	5
23	The association between sunspot magnetic fields and superpenumbral fibrils. Astronomische Nachrichten, 2014, 335, 161-167.	1.2	3
24	NLTE modeling of a small active region filament observed with the VTT. Astronomische Nachrichten, 2016, 337, 1045-1049.	1.2	3
25	Spectral Characteristics of the He i D3 Line in a Quiescent Prominence Observed by THEMIS. Solar Physics, 2017, 292, 1.	2.5	2
26	Magnetic Flux Emergence in a Coronal Hole. Solar Physics, 2020, 295, 1.	2.5	2
27	Tracking Downflows from the Chromosphere to the Photosphere in a Solar Arch Filament System. Astrophysical Journal, 2020, 890, 82.	4.5	1
28	The three-dimensional structure of the magnetic field of a sunspot. Proceedings of the International Astronomical Union, 2008, 4, 225-226.	0.0	0
29	Transmission profile of the Dutch Open Telescope Hα Lyot filter. Astronomische Nachrichten, 2014, 335, 409-416.	1.2	0
30	Flows along arch filaments observed in the GRIS †very fast spectroscopic mode'. Proceedings of the International Astronomical Union, 2016, 12, 28-33.	0.0	0