

Victor Yanke

List of Publications by Year in descending order

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66
papers

934
citations

567281

15
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477307

29
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69
all docs

69
docs citations

69
times ranked

474
citing authors

#	ARTICLE	IF	CITATIONS
1	Space weather conditions and spacecraft anomalies in different orbits. <i>Space Weather</i> , 2005, 3, n/a-n/a.	3.7	116
2	Magnetospheric effects in cosmic rays during the unique magnetic storm on November 2003. <i>Journal of Geophysical Research</i> , 2005, 110, .	3.3	101
3	Coronal Mass Ejections and Non-recurrent Forbush Decreases. <i>Solar Physics</i> , 2014, 289, 3949-3960.	2.5	74
4	The Global Survey Method Applied to Ground-level Cosmic Ray Measurements. <i>Solar Physics</i> , 2018, 293, 1.	2.5	54
5	Global Survey Method for the World Network of Neutron Monitors. <i>Geomagnetism and Aeronomy</i> , 2018, 58, 356-372.	0.8	52
6	Galactic Cosmic Ray Density Variations in Magnetic Clouds. <i>Solar Physics</i> , 2015, 290, 1429-1444.	2.5	49
7	Ground level enhancements of solar cosmic rays during the last three solar cycles. <i>Geomagnetism and Aeronomy</i> , 2010, 50, 21-33.	0.8	45
8	Precursor Effects in Different Cases of Forbush Decreases. <i>Solar Physics</i> , 2012, 276, 337-350.	2.5	35
9	Forbush effects with a sudden and gradual onset. <i>Geomagnetism and Aeronomy</i> , 2012, 52, 292-299.	0.8	33
10	Main Properties of Forbush Effects Related to High-Speed Streams from Coronal Holes. <i>Geomagnetism and Aeronomy</i> , 2018, 58, 154-168.	0.8	30
11	Solar cosmic rays during the extremely high ground level enhancement on 23 February 1956. <i>Annales Geophysicae</i> , 2005, 23, 2281-2291.	1.6	26
12	Temperature effect of the muon component and practical questions for considering it in real time. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2011, 75, 820-824.	0.6	21
13	The Asymptotic Longitudinal Cosmic Ray Intensity Distribution as a Precursor of Forbush Decreases. <i>Solar Physics</i> , 2012, 280, 641-650.	2.5	18
14	Forbush Decreases Associated with Western Solar Sources and Geomagnetic Storms: A Study on Precursors. <i>Solar Physics</i> , 2013, 283, 557-563.	2.5	17
15	Relationship between Forbush effect parameters and the heliolongitude of solar sources. <i>Geomagnetism and Aeronomy</i> , 2013, 53, 10-18.	0.8	15
16	Ring of Stations Method in Cosmic Rays Variations Research. <i>Solar Physics</i> , 2020, 295, 1.	2.5	15
17	Efficiency of detection for neutron detectors with different geometries. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2011, 75, 866-868.	0.6	14
18	Behavior of the cosmic-ray vector anisotropy before interplanetary shocks. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2009, 73, 331-333.	0.6	12

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19	Magnetospheric effects of cosmic rays. 1. Long-term changes in the geomagnetic cutoff rigidities for the stations of the global network of neutron monitors. <i>Geomagnetism and Aeronomy</i> , 2016, 56, 381-392.	0.8	10
20	Long-Term Changes in the Number and Magnitude of Forbush-Effects. <i>Geomagnetism and Aeronomy</i> , 2018, 58, 615-624.	0.8	10
21	Procedure to emend neutron monitor data that are affected by snow accumulations on and around the detector housing. <i>Journal of Geophysical Research: Space Physics</i> , 2013, 118, 6852-6857.	2.4	9
22	Estimation of long-term stability of detectors within the global network of neutron monitors. <i>Geomagnetism and Aeronomy</i> , 2007, 47, 251-255.	0.8	8
23	Recurrent and sporadic Forbush-effects in deep solar minimum. <i>Journal of Physics: Conference Series</i> , 2015, 632, 012062.	0.4	8
24	Possible ground level enhancements of solar cosmic rays in 2012. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2015, 79, 561-565.	0.6	8
25	Cosmic-ray vector anisotropy and local characteristics of the interplanetary medium. <i>Geomagnetism and Aeronomy</i> , 2017, 57, 389-397.	0.8	8
26	Solar wind temperatureâ€“velocity relationship over the last five solar cycles and Forbush decreases associated with different types of interplanetary disturbance. <i>Monthly Notices of the Royal Astronomical Society</i> , 2020, 500, 2786-2797.	4.4	8
27	Simulation of the modulation of galactic cosmic rays during solar activity cycles 21â€“23. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2007, 71, 974-976.	0.6	7
28	Long-period variations in the amplitude-phase interrelation of the first cosmic ray anisotropy harmonic. <i>Geomagnetism and Aeronomy</i> , 2013, 53, 561-570.	0.8	7
29	Forbush-decreases in 19th solar cycle. <i>Journal of Physics: Conference Series</i> , 2013, 409, 012165.	0.4	7
30	Coronal holes in the long-term modulation of cosmic rays. <i>Geomagnetism and Aeronomy</i> , 2016, 56, 257-263.	0.8	7
31	Contributions from changes in various solar indices in cycles 20â€“23 and 24 to the modulation of cosmic rays. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2017, 81, 146-150.	0.6	7
32	Forbush Effects Created by Coronal Mass Ejections with Magnetic Clouds. <i>Geomagnetism and Aeronomy</i> , 2021, 61, 678-687.	0.8	7
33	Behavior of the Speed and Temperature of the Solar Wind during Interplanetary Disturbances Creating Forbush Decreases. <i>Geomagnetism and Aeronomy</i> , 2020, 60, 521-529.	0.8	7
34	Effect of snow in cosmic ray variations and methods for taking it into consideration. <i>Geomagnetism and Aeronomy</i> , 2011, 51, 247-253.	0.8	6
35	Behavior of the cosmic ray density during the initial phase of the Forbush effect. <i>Geomagnetism and Aeronomy</i> , 2016, 56, 645-651.	0.8	6
36	Vector anisotropy of cosmic rays in the beginning of Forbush effects. <i>Geomagnetism and Aeronomy</i> , 2017, 57, 541-548.	0.8	6

#	ARTICLE	IF	CITATIONS
37	Influence of Interacting Solar Wind Disturbances on the Variations in Galactic Cosmic Rays. <i>Geomagnetism and Aeronomy</i> , 2021, 61, 792-800.	0.8	6
38	Long-Term Changes in Vertical Geomagnetic Cutoff Rigidities of Cosmic Rays. <i>Physics of Atomic Nuclei</i> , 2018, 81, 1382-1389.	0.4	5
39	Spectrum of long-term cosmic ray variations during the sunspot minimum in 2009. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2013, 77, 513-516.	0.6	4
40	Analyzing the temperature effect of high mountain cosmic ray detectors using the database of the global network of muon telescopes. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2015, 79, 662-666.	0.6	4
41	Temperature Effect Observed for the Muon Component in the Yakutsk Cosmic-Ray Spectrograph. <i>Physics of Atomic Nuclei</i> , 2018, 81, 776-785.	0.4	4
42	Method of Global Survey by Data of Muon Telescopes. <i>Physics of Atomic Nuclei</i> , 2019, 82, 879-885.	0.4	4
43	Features of the Behavior of Time Parameters of Forbush Decreases Associated with Different Types of Solar and Interplanetary Sources. <i>Geomagnetism and Aeronomy</i> , 2022, 62, 17-31.	0.8	4
44	Size Distributions of Solar Proton Events: Methodological and Physical Restrictions. <i>Solar Physics</i> , 2016, 291, 3685-3704.	2.5	3
45	Planetary long term changes of the cosmic ray geomagnetic cut off rigidities. <i>Journal of Physics: Conference Series</i> , 2019, 1181, 012008.	0.4	3
46	Large Forbush Decreases and their Solar Sources: Features and Characteristics. <i>Solar Physics</i> , 2020, 295, 1.	2.5	3
47	Forbush decreases caused by paired interacting solar wind disturbances. <i>Monthly Notices of the Royal Astronomical Society</i> , 2022, 511, 5897-5908.	4.4	3
48	Relationship between Forbush effects and X-ray flares. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2007, 71, 988-990.	0.6	2
49	Cosmic ray events in the beginning of 2012. <i>Journal of Physics: Conference Series</i> , 2013, 409, 012206.	0.4	2
50	Density variations of galactic cosmic rays in magnetic clouds. <i>Geomagnetism and Aeronomy</i> , 2015, 55, 430-441.	0.8	2
51	Distribution of Cosmic Rays in the Heliosphere, According to Data from the Network of Muon Telescope Stations. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2019, 83, 547-550.	0.6	2
52	Precursory Signs of Large Forbush Decreases. <i>Solar Physics</i> , 2021, 296, 1.	2.5	2
53	Long-Term Modulation of Cosmic Rays in Solar Cycles 23â€“24. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2021, 85, 1045-1048.	0.6	2
54	Using Data from a Ground-Based Network of Detectors and the PAMELA and AMS-02 Experiments to Compare Long-Term Variations in the Cosmic Ray Flux. <i>Bulletin of the Russian Academy of Sciences: Physics</i> , 2021, 85, 1039-1041.	0.6	2

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55	Program for Calculating the Geomagnetic Cutoff Rigidity of Cosmic Rays and the Trajectories of Their Motion. Bulletin of the Russian Academy of Sciences: Physics, 2021, 85, 1297-1301.	0.6	2
56	Forbush decreases in the 19th solar cycle. Bulletin of the Russian Academy of Sciences: Physics, 2013, 77, 535-537.	0.6	1
57	Seasonal variations of the muon flux in the MUSTANG super telescope data. Bulletin of the Russian Academy of Sciences: Physics, 2013, 77, 561-565.	0.6	1
58	Modeling variations in CR density in magnetic clouds. Bulletin of the Russian Academy of Sciences: Physics, 2015, 79, 637-639.	0.6	1
59	Size Distribution of Forbush Effects. Geomagnetism and Aeronomy, 2018, 58, 809-816.	0.8	1
60	Experimental Spectrum of Cosmic Ray Variations in Earth Orbit, According to AMS-02 Data. Bulletin of the Russian Academy of Sciences: Physics, 2021, 85, 1042-1044.	0.6	1
61	Recognition of Geomagnetic Storm Based on Neural Network Model Estimates of Dst Indices. Journal of Computer and Systems Sciences International, 2022, 61, 54-64.	0.6	1
62	Large Scale Modulation: View from the Earth Points. Physics of Atomic Nuclei, 2021, 84, 1159-1170.	0.4	1
63	Anomalous Forbush effects from sources far from Sun center. Proceedings of the International Astronomical Union, 2008, 4, 451-456.	0.0	0
64	Long-Term Trends in Forbush Decrease Activity over the Last Six Solar Cycles. Bulletin of the Russian Academy of Sciences: Physics, 2019, 83, 566-568.	0.6	0
65	An Unusual Decrease in the Intensity of Cosmic Rays in May 2019 during the Solar Minimum. Bulletin of the Russian Academy of Sciences: Physics, 2021, 85, 588-591.	0.6	0
66	Heliospheric Modulation of Cosmic Rays in the Era of Neutron Monitoring. Bulletin of the Russian Academy of Sciences: Physics, 2021, 85, 1052-1054.	0.6	0